

CONSTELLATION X-RAY OBSERVATORY

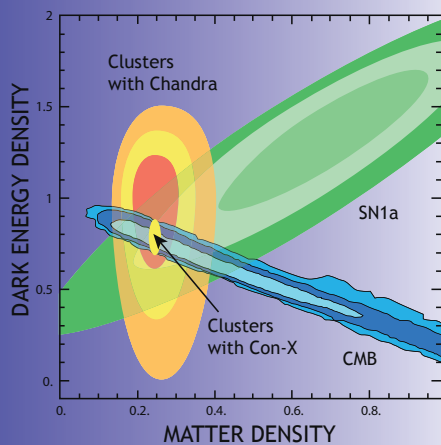
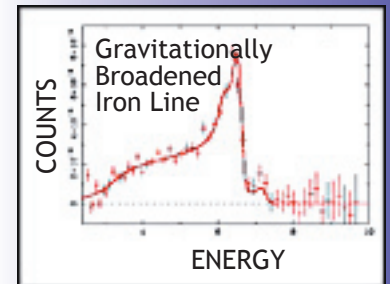
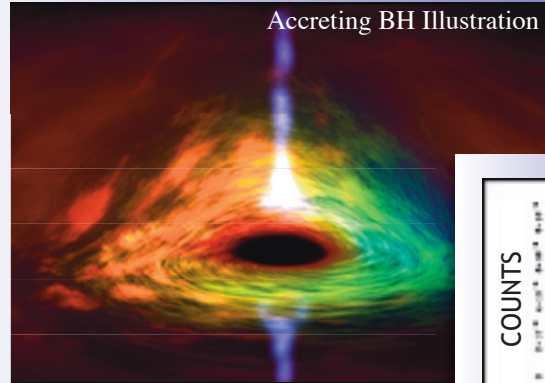
A BEYOND EINSTEIN GREAT OBSERVATORY

TO UNDERSTAND THE GREAT MYSTERIES OF SPACE, TIME AND ENERGY

What happens to space, time, and matter at the edge of a black hole?

- Test limits of General Relativity in extreme gravity at black hole event horizon
- Measure spin of black holes
- Make first “movies” of matter spiraling into a black hole

B L A C K H O L E S



D A R K E N E R G Y

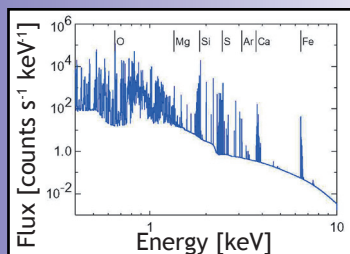
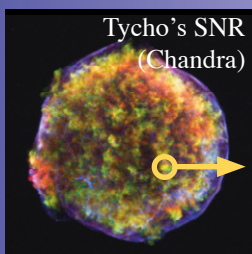
What is the Universe made of?

- Observe clusters and growth of structure to measure expansion of Universe and determine properties of Dark Energy (DE)
- Measure Dark Matter (DM) content of galaxies, groups, and clusters.
- X-ray observations of clusters see the 4% normal matter, and constrain the nature of the 26% DM and 70% DE.

How did the Universe come to look like it does now?

- Understand connection between growth of supermassive black holes and host galaxies. Self-regulating accretion alternating with outflows from nuclear black holes?
- Study superwind feedback mechanism in starburst galaxies

C O S M I C F E E D B A C K



L I F E C Y C L E S
O F M A T T E R

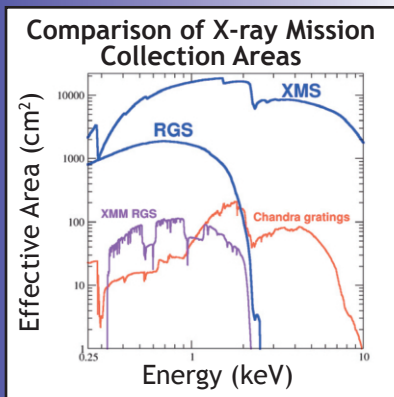
What is the origin of atoms in stars, planets, and living organisms?

- Trace formation of individual elements in SN explosions
- Probe formation of stars and planets
- Determine nature of superdense matter in neutron stars

The 2000 NRC Decadal Survey (Astronomy and Astrophysics in the New Millennium) **ranked Con-X next in priority after JWST** among major space-based initiatives. These Decadal Survey priorities were re-affirmed by a 2005 NRC Mid-Course Review. Con-X has also received strong endorsement from the Quarks to Cosmos NRC report.



CONSTELLATION-X: MISSION AND TECHNOLOGY



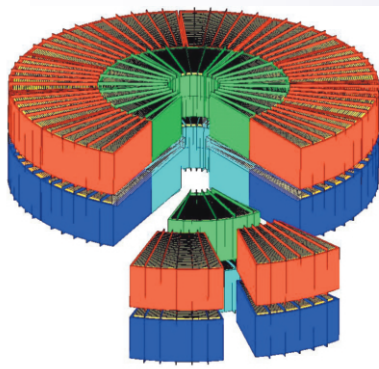
MISSION OVERVIEW

- L2 orbit for high viewing efficiency and stable thermal environment
- 5 year lifetime with 10 year goal
- Technically ready, well understood, mission with simple spacecraft

Con-X effective area vs energy compared to Chandra and XMM spectrometers. 100-fold increase in collecting area provides for breakthrough science.

KEY REQUIREMENTS

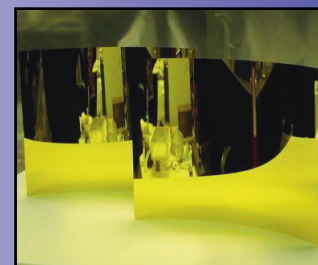
Bandpass: 0.25 to 40keV
Area: 15000cm²@1.25keV,
 6000cm²@6 keV,
 1500cm²@40 keV
Spatial Resolution:
 15" HPD, 5" goal, for SXT
 1' HPD, 30" goal, for HXT
Spectral Revolving Power:
 >300 from 0.25keV to 6keV
 1500 from 6keV to 10keV



MIRRORS

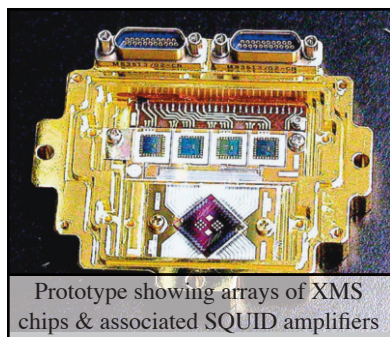
- Spectroscopy X-ray Telescope (SXT) mirror technologies derive directly from flight programs (XMM-Newton, Suzaku) but with improved figure accuracy and reduced mass.
- Assembled from 60 (30) degree wedges into circular mirror. Reflection gratings behind SXT.
- Con-X uses thermally slumped glass coated with gold.
- Hard X-ray Telescope (HXT) mirrors use similar approach with multi-layer coatings to extend to higher energies.

Left: A thermally formed glass segment being lifted off a mandrel.
Right: Two slumped glass segments coated with the gold reflective surface.

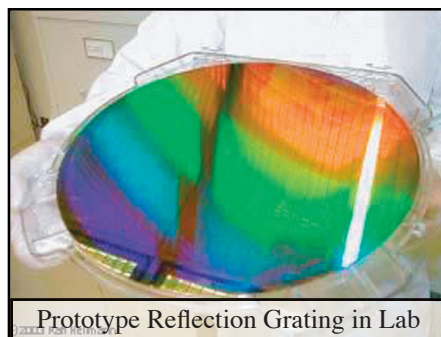


SCIENCE INSTRUMENTS

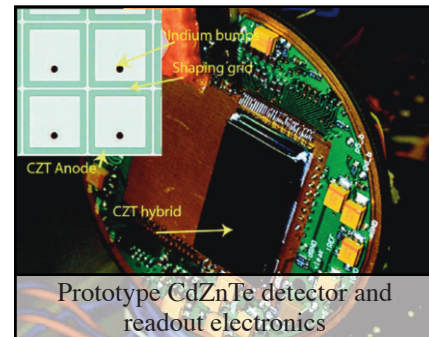
X-ray Microcalorimeter Spectrometer (XMS). Provides imaging, non-dispersive, high resolution spectroscopy at SXT focus. Superconducting device, closed loop cryo-cooler, no stored (expendable) cryogen.



Reflection Grating Spectrometer (RGS). Provides high resolution, dispersive spectroscopy at lower energies. Attached directly behind part of SXT. Read-out with CCDs



Hard X-ray Telescope (HXT) mirrors uses segments or full shells with multi-layer coatings to extend to higher energies. Provides imaging, moderate resolution spectroscopy at highest energies using hybrid CdZnTe arrays.



X-ray spectroscopy now rivals the optical for breadth and depth of science. The technologies needed for Con-X are well understood and performance has been demonstrated.

<http://constellation.gsfc.nasa.gov>